




#556

ISEE 1

PLASMASPHERIC ELEC. FLD. 3 SEC. AVG

77-102A-11D



ISEE 1

3-S AVGD PLASMASPHERIC ELEC. FIELD

77-102A-11D

THIS DATA SET HAS BEEN RESTORED. THERE WAS ORIGINALLY ONE 9-TRACK, 1600 BPI TAPE WRITTEN IN EBCDIC. THERE IS ONE RESTORED TAPE WRITTEN IN ASCII. THE DR TAPE IS A 3480 CARTRIDGE AND THE DS TAPE IS 9-TRACK, 6250 BPI. THE TIME SPAN IS NOT CONTAINED ON THE TAPE AND CAN NOT BE VERIFIED. THE ORIGINAL TAPE WAS CREATED ON AN IBM 360 COMPUTER AND WAS RESTORED ON AN IBM 9021 COMPUTER. THE DR AND DS NUMBER ALONG WITH THE CORRESPONDING D NUMBER AND TIME SPAN IS AS FOLLOWS:

DR#	DS#	D#	FILES	TIME SPAN
DR005346	DS005346	D053321	160	12/02/77 - 11/30/78

REQ. AGENT

DEW

REQ. NO.

V0160

ACQ. AGENT

HKH

ISEE 1
PLASMASPHERIC ELEC. FLD. 3 SEC. AVG
77-102A-11D

This data set catalog consists of 1 data tape. The tape is 1600, bpi, 9 track, EBCDIC, with 160 files of data. The time span is not on the tape. The D and C numbers are as follows:

<u>D#</u>	<u>C#</u>	<u>TIME SPAN</u>	<u>FILES</u>
D-53321	C-22722	12/2/77 - 11/30/78	160

September 15, 1982

TO: 601/NSSDC/Kent Hills
FROM: 696/Computer Sciences Corporation
SUBJECT: 9 Track 1600 bpi Tape of the ISEE-1

Attached is a 9 track 1600 bpi tape of the ISEE-1 culled data sets for a 12 month period beginning in December 1977. In addition, there is a chronological listing of the files noting the start and finish times and whether the satellite was in-bound or out-bound. There are 166 files on the tape and the following is a list of their characteristics:

The recordlength is 55 bytes (the 55th byte is a control character) and the block size is 2750 bytes.

Format (3I2, 6F8.3, 1x)

	TIME								
	HH	MM	SS	L	MLT	E _n	E _E	E _R	E _A
Byte#	12, 34, 56,	7..14,	15..22,	23..30,	31..38,	39..46,	47..54		

where E_n and E_E are the electric field components in the equatorial plane and E_R and E_A are their ionospheric projections. A data description is attached.

Please contact me if you have any questions.

Cathie Meetre

Cathie Meetre (344-5403)

Attachments: 2

September 14, 1982

TO: 601/NSSDC/Kent Hills
FROM: 696/Electrodynamics Branch
SUBJECT: ISEE-1 Electric Field Data

Attached is (1) a data tape, (2) tape format, (3) listing of files and (4) a brief description of how the data was handled for ISEE-1 Electric Field data from the HPM instrument. The tape covers from December 1, 1977 to December 1, 1978 and includes all usable data in the plasmasphere (mostly inside L of 6). The listing of files gives each UT time span. We would like to submit this data to the NSSDC ISEE files. For more information contact N. C. Maynard.



James P. Heppner

Attachments: 4

HPM PLASMAPHERIC ELECTRIC FIELDS

The Electric Field Data

Electric field measurements were made from the ISEE-1 spacecraft using the double probe technique with long cylindrical sensors (Heppner et al., 1978). The baseline (separation between the midpoints of the sensors) of the double probe system of long wires was 179 m. From this single axis, measurements of both components of the electric field in the spin plane were possible every spin period (3.05s). Data along the instantaneous direction of the axis were taken 16 or 32 times per second depending on the data rate. The spin plane was approximately coincident with the solar ecliptic XY plane in this 23.2 Re apogee orbit.

In all data presented here the $\underline{v}_s \times \underline{B}$ electric field from the satellite velocity \underline{v}_s and the corotation electric field have been vectorially subtracted. Thus the measurements are transferred to a coordinate system corotating with an observer on the Earth. The magnetic field used in the $\underline{v}_s \times \underline{B}$ calculations was that measured on board the spacecraft (Russell, 1978).

To determine the electric field in the spin plane, data taken when either antenna was within 20° of the sun was culled in order to remove any errors caused by shadowing of a sensor or by reduced photoemission from a small sun angle, and the remaining data was fitted to a sine wave using a least squares technique. The solar ecliptic X and Y electric field components in inertial space were then determined from the sine wave fit, and the $\underline{v}_s \times \underline{B}$ and corotation subtractions were done to yield the electric field in the earth fixed coordinate system.

The third component of the electric field (solar ecliptic E_z) was calculated assuming that $\underline{E} \cdot \underline{B} = 0$. This extrapolation is good providing that the ratio of \underline{B}_{xy} to \underline{B}_z is low. The resulting vector electric field was then transformed into the radially outward and azimuthally eastward

components perpendicular to \underline{B} . For the plasmasphere studies these electric fields were then projected to the magnetic equatorial plane and to the ionosphere assuming that the magnetic field lines are equipotentials and that the magnetic field was a dipole. The projections were done on each data set to provide a common reference for comparison of data taken at different magnetic latitudes. When projected to the northern hemisphere ionosphere these components become the northward and eastward electric fields respectively. Note that the projection introduces errors from our imprecise knowledge of the magnetic field configuration. The original L values were calculated using the GSFC harmonic analyses of B with the Mead-Fairfield tail extension. We in turn have projected the values to the ionosphere assuming a symmetric dipole. Thus especially at the higher L values, both the projection factor and location will have some error. This was not considered to be critical to the results presented here.

The purpose of this study was to find the average electric field seen in the plasmasphere. Thus it was mandatory to remove any bad data that could affect the results. All data were scanned, and regions where the instrument was in its internal calibration cycle as well as regions where the Harvey instrument (Harvey et al., 1978) was transmitting (causing the electric field sensors to be shorted) were culled from the data set.

In regions of low density plasma ($< 1/\text{cm}^3$) the electric field instrument sees an apparent sunward directed electric field that is caused by the asymmetries of the sheath of photoelectrons around the spacecraft. The effect results from two causes. The dipole of the negatively charged sheath on the sunward side of the spacecraft and the positively charged spacecraft is one cause, and the other is the asymmetric collection by the sunward sensor of photoelectrons from the spacecraft, (see Cauffman and Maynard, 1975). It is hard to pinpoint exactly at what density this effect becomes significant (it is also a function of the temperature of the medium as well). As the plasma becomes more rarified, the first sign of asymmetrical photoelectron emission errors appears in data taken with the sensor axis within $\pm 20^\circ$ from the sun or at a near grazing angle (data which is not used in the least squares calculation to determine the electric field). In order to scrupulously avoid any data that might be the least

bit contaminated by this effect, all data were culled that had large bite outs where the antenna axis was directed toward the sun or that had a significant solar ecliptic E_x . This last criteria eliminated some good data where the convective electric field would be expected to be in that direction. A comparison of the point where culling started and plots of electron density from the Harvey instrument for several orbits (M. Echeto, private communication, 1980) showed that the data that were kept were in general from regions of electron densities greater than 30 to 50/cm³.

The ISEE-1 electric field data from 1 December 1977 to 1 December 1978 were scanned and culled with the above criteria to form the data base for this study. A second culling of the data base was done to remove data when the ratio of B_z to $|B|$ was 0.2 or less in order to prevent errors from the $\underline{E} \cdot \underline{B}$ extrapolation from seriously effecting the results.

ISFF-1 Culled data sets
in chronological order
December 1977 - November 1978

Prepared by C. Heetie, C.S.C.
July 1982

December 77

Start time Stop time Tape File # I O N Disc #

Date

DEC 2, 77	4	54	50	5	30	0	1	✓			✓	14
DEC 4, 77	13	45	02	14	56	26	2	✓			✓	14
DEC 4, 77	15	58	25	16	29	55	3		✓		✓	14
DEC 6, 77	23	0	1	23	59	55	4	✓			✓	14
DEC 7, 77	0	0	7	0	19	44	5	✓			✓	14
DEC 7, 77	1	24	26	2	04	55	6		✓		✓	14
DEC 9, 77	10	48	35	11	08	38	7		✓		✓	14
DEC 9, 77	7	40	1	9	9	16	8	✓			✓	14
DEC 11, 77	18	12	06	19	09	59	9	✓			✓	14
DEC 14, 77	3	42	40	4	24	59	10	✓			✓	14
DEC 14, 77	5	32	12	5	37	49	11		✓		✓	14
DEC 16, 77	14	55	28	15	30	59	12		✓		✓	1
DEC 16, 77	11	38	36	13	34	57	13	✓			✓	1
DEC 18, 77	21	48	06	23	09	10	14	✓			✓	1
DEC 19, 77	0	24	06	0	54	58	15		✓		✓	1
DEC 21, 77	9	44	44	9	57	53	16	✓			✓	1
DEC 21, 77	7	40	06	8	14	59	17	✓			✓	1
DEC 23, 77	19	13	07	19	33	06	18		✓		✓	1
DEC 23, 77	16	04	07	17	59	03	19	✓			✓	1
DEC 24, 77	2	13	7	3	13	9	20	✓			✓	2
DEC 28, 77	13	52	03	14	19	55	21		✓		✓	1
DEC 28, 77	10	41	7	12	40	55	22	✓			✓	1
DEC 30, 77	20	22	06	21	51	50	23	✓			✓	2
DEC 30, 77	23	14	38	23	29	58	24		✓		✓	2
mislabelled on disc as DEC 12 78 IC												

(33)

Date	Start time		Stop time		Tape file #	I	O	N	Disc #
	Hr	Min	Hr	Min					
Jan 2, 1978	6	32	00	00	25	✓			✓
Jan 2	8	40	43	59	26		✓		✓
Jan 4	18	8	16	18	27		✓		✓
Jan 4	16	19	02	43	28	✓			✓
Jan 9	12	52	11	59	29		✓		✓
Jan 9	10	22	26	28	30	✓			✓
Jan 11	19	38	01	49	31	✓			✓
Jan 11	22	12	14	18	32		✓		✓
Jan 14	4	57	41	15	33	✓			✓
Jan 16	13	20	03	24	34	✓			✓
Jan 18	23	49	0	59	35	✓			✓
Jan 19	0	0	9	46	36	✓			✓
Jan 21	10	59	18	1	38	✓			✓
Jan 21	8	15	1	7	37		✓		✓
Jan 23	21	29	50	37	39		✓		✓
Jan 23	17	15	02	19	40	✓			✓
Jan 26	5	39	08	52	41		✓		✓
Jan 28	15	4	2	59	42		✓		✓
Jan 28	12	38	02	20	43	✓			✓
Jan 30	22	55	1	13	44	✓			✓
Jan 31	20	27	1	53	45		✓		✓

Date	Start time			Stop time			Tape (ch)	In	Out	N	Disc#
	HR	Min	Sec	HR	Min	Sec					
Feb 02, 1978	9	50	08	10	39	59	46	✓	✓		15
Feb 02	18	35	1	9	6	22	59	✓	✓		4
Feb 04	19	9	7	19	54	55	47	✓	✓		15
Feb 07	DID	NOT	NOT	COPY			48	✓	✓		16
Feb 09	12	20	2	12	54	58	48	✓	✓		4
Feb 11	21	25	1	22	29	58	49	✓	✓		4
Feb 14	6	28	44	7	54	59		✓	✓		4
Feb 14	8	40	8	9	27	48		✓	✓		4
Feb 16	18	7	40	18	42	57	50	✓	✓		16
Feb 16	16	43	10	17	19	59	51	✓	✓		16
Feb 19	3	23	8	3	44	57	52	✓	✓		4
Feb 21	12	44	7	13	34	58	53	✓	✓		16
Feb 21	10	49	33	12	4	54	54	✓	✓		4
Feb 23	22	7	16	22	44	57	55	✓	✓		16
Feb 23	DID	NOT	NOT	COPY			56	✓	✓		4
Feb 26	7	32	8	8	9	58	57	✓	✓		16
Feb 26	5	30	1	6	45	0	58	✓	✓		3
Feb 28	15	38	2	16	15	0		✓	✓		4
Feb 28	16	57	6	17	34	54	60	✓	✓		3
							61	✓	✓		16

Date archi '78	Start time				Type file #	In Out		Disc #
	HR	MIN	SEC	HR		MIN	SEC	
3	2	17	7	2	62	✓	56	10
5	9	39	6	10	63	✓	59	10
5	11	39	48	12	64	✓	0	10
7	19	40	7	20	65	✓	58	10
10	4	30	8	5	66	✓	18	10
10	6	26	8	6	67	✓	58	10
12	13	53	7	15	68	✓	58	17
12	15	49	7	16	69	✓	58	17
15	0	10	8	2	70	✓	9	17
17	10	34	7	11	71	✓	26	17
17	9	8	8	9	72	✓	59	17
29	7	42	7	8	72	✓	57	5
29	9	48	7	10	73	✓	58	14
31	18	46	6	19	74	✓	58	14

May, 78

Date _____
Page 78

start

stop

Tape file #

208

2

Jan 18	23 DEC 16	23	29	37	0 N944				
8	8	8	9	12	20	92	✓	✓	8
11	22	9	12	20	92	✓	✓	✓	8
13	TRX THIS AGAIN						✓	✓	4
16	3 24	3	53	52	93	✓	✓	✓	9
16	4 35	5	32	59	94	✓	✓	✓	9
18	13 3	13	10	0	95	✓	✓	✓	9
23	8 0 25	8	8	1	96	✓	✓	✓	9
25	8 56	9	19	43	97	✓	✓	✓	9
28	2 30	2	42	50	98	✓	✓	✓	9
28	1 5	1	22	32	99	✓	✓	✓	9
28	3 22	4	16	36	100	✓	✓	✓	9

June '78

In Out N

Tape file #

Stop

Start

Date

June '78

1	22	15	25	22	34	58	101	✓	12
1	20	0	6	20	18	46	102	✓	12
4	5	30	4	25	49	58	103	✓	12
4	7	36	9	8	18	6	104	✓	12
9	2	6	9	2	39	57	105	✓	12
11	11	25	8	12	24	59	106	✓	10
11	9	43	22	9	53	8	107	✓	10
13	20	52	6	20	57	29	108	✓	10
16	5	28	39	5	34	59	109	✓	10
20	22	41	4	23	25	0	110	✓	10
21	0	57	9	1	42	43	111	✓	10
23	8	31	8	8	51	58	112	✓	10
23	9	30	40	9	36	48	113	✓	10
23	10	9	8	10	39	45	114	✓	10
28	3	4	9	3	37	18	115	✓	10

July 78

Disc #

N

O

I

Tape file #

Stop

Start

Date

July 78

2	22	4	36	23	23	55	116	✓				16
2	23	34	7	23	59	58	117		✓			16
3	20	50	19	0	10	0	118		✓			16
5	9	2	15	9	29	58	119		✓			16
10	1	36	17	2	28	32	120	✓				16
12	10	56	18	11	53	24	121	✓				16
14	21	0	6	21	19	59	122	✓				11
17	5	0	7	5	10	28	123	✓				15
17	8	22	16	8	53	52	124		✓			11
19	15	20	9	15	53	52	125	✓				11
22	0	40	9	1	21	26	126	✓				11
22	2	16	6	3	29	27	127		✓			11
24	9	54	7	10	43	56	128					11
26	19	0	6	19	14	58	129		✓			11
26	20	58	0	21	45	0	130		✓			11
29	7	8	9	7	39	57	131		✓			11
29	4	25	2	5	29	57	132	✓				11

Disc#

I O W

Tape file #

Step

Start

Date

Aug '78

3	1	21	6	2	8	19	133	✓		15
4							132			
5							131	✓	←	15
10	5	6	7	6	4	58	134	✓		15
10	3	15	1	4	20	0	135	✓		15
14	23	47	9	23	54	58	136	✓		15
15	0	11	18	0	24	59	137	✓		15
17	7	8	8	7	34	58	138	✓		15
18							137	✓		15
19	18	27	08	18	36	37	139	✓		15
22	3	51	31	4	9	59	140	✓		15
22	1	4	8	3	9	59	141	✓		15

\$NOP

\$NOP ***** DO-1 *****

\$EXEC TPLIST BS

053321

12/2/77 - 11/30/78

INPUT PARAMETERS ARE: ED SR=1=1 1

TAPE NO.	1	FILE NO.	1
RECORD	1	LENGTH	2750
45450	4.211	11.998	-1.615 -6.424 -0.103 -0.744 45453 4.209 11.998 -1.175 -8.526 -0.0
75	-3.987	45456	4.208 11.999 -0.357 -8.769 -0.023 -1.016 45459 4.206 12.001 -0.929 -
7.372	-3.059	-0.835	455 2 4.204 12.002 -1.516 -0.097 -0.818 455 5 4.202 12.003
-2.365	-8.592	-3.151	-0.927 455 8 4.201 12.005 -4.494 -9.362 -0.268 -1.087 45511 4.
199	12.006	-5.633	-7.372 -0.361 -0.857 45514 4.197 12.008 -5.523 -7.797 -0.354 -0.907
45517	4.196	12.009	-5.427 -8.740 -0.348 -1.040 45520 4.194 12.010 -6.145 -9.902 -0.3
95	-1.153	45523	4.192 12.012 -6.272 -10.529 -0.403 -1.227 45526 4.190 12.013 -3.707 -1
0.622	-1.239	-1.233	45529 4.189 12.015 -1.719 -9.983 -0.111 -1.164 45532 4.187 12.016
-2.693	-6.390	-1.173	-0.735 45535 4.185 12.017 -3.280 -6.554 -0.211 -0.766 45538 4.
183	12.019	-5.230	-5.328 -0.337 -0.623 45541 4.182 12.020 -6.425 -1.579 -0.415 -0.185
45544	4.180	12.022	-5.792 0.457 -0.374 0.053 45547 4.178 12.023 -6.723 0.142 -0.4
35	0.917	45553	4.176 12.024 -8.317 0.919 -0.538 0.108 45554 4.175 12.026 -5.659
0.524	-0.366	0.661	45557 4.173 12.027 -2.716 -2.289 -0.176 -0.269 456 0 4.171 12.029
-3.617	-1.328	-0.234	-0.156 456 3 4.174 12.030 -1.448 2.112 -0.004 0.248 456.6 4.
168	12.031	-0.537	5.952 -0.035 0.699 456 9 4.166 12.033 1.782 6.197 0.116 0.729
45612	4.164	12.034	2.482 6.370 0.161 0.750 45615 4.163 12.035 2.340 9.257 0.1
52	1.000	45618	4.161 12.037 1.096 9.866 0.071 1.162 45621 4.159 12.039 3.199 1
0.745	0.208	1.267	45624 4.157 12.040 -4.262 7.210 -0.278 0.851 45627 4.156 12.041
-8.157	6.883	-1.525	0.513 45630 4.154 12.043 -8.007 8.214 -0.522 0.970 45633 4.
152	12.044	-12.121	8.538 -0.791 1.099 45636 4.151 12.046 -12.512 10.305 -0.817 1.218
45639	4.149	12.047	-12.581 11.351 -0.822 1.343 45642 4.147 12.049 -14.734 11.445 -0.9
64	1.355	45645	4.145 12.050 -15.023 19.826 -0.987 1.283 45648 4.144 12.051 -14.506 1
0.656	-0.951	1.263	45651 4.142 12.053 -12.016 9.489 -0.788 1.126 45654 4.140 12.054
-6.708	7.284	-0.440	0.864 45657 4.138 12.056 -4.883 6.909 -0.321 0.821 457 0 4.
137	12.057	-5.911	6.747 -0.388 0.802 457 3 4.135 12.059 -6.318 7.734 -0.415 0.920
457 6	4.133	12.060	-7.204 7.736 -0.520 0.921 45710 4.132 12.062 -9.480 5.179 -0.6
24	0.617	45713	4.131 12.063 -9.727 4.427 -0.641 0.527 45716 4.128 12.065 -10.653
3.731	-0.702	0.451	45719 4.126 12.066 -10.935 2.274 -0.721 0.271

***** JOB DONE.

\$WE0 LPS

\$s

\$ASS IN ID2 OUT MT4

\$EXE TPOUPC BS